

Degeneracies between Overlapping and Lensing GWs

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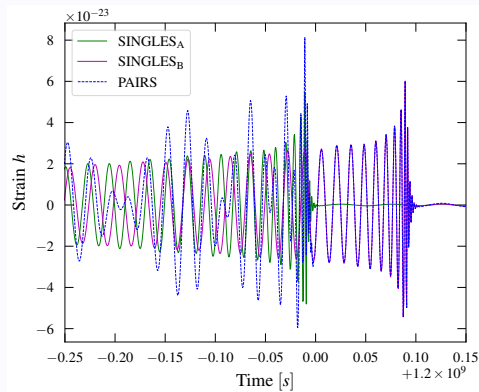
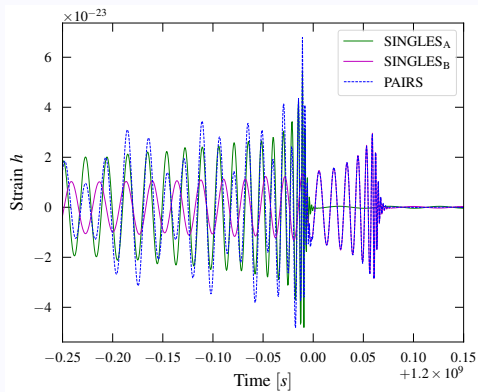
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Agenda

- 1 Overview
- 2 Introduction
- 3 Lensing Effects
- 4 Analysis Setup
- 5 Results
- 6 Conclusions

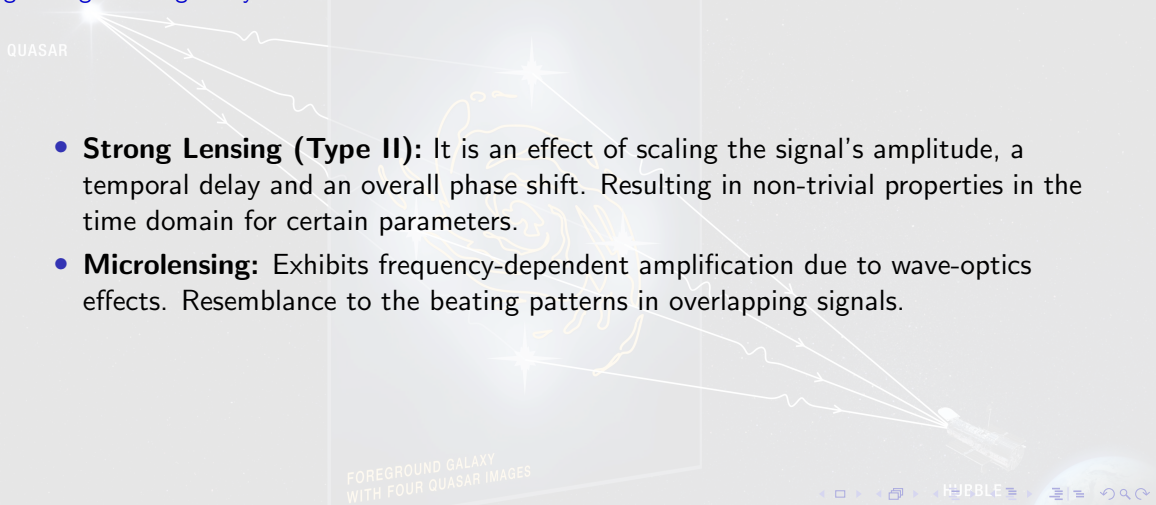
Overlapping Transients



Overlapping Gravitational Wave Signals.

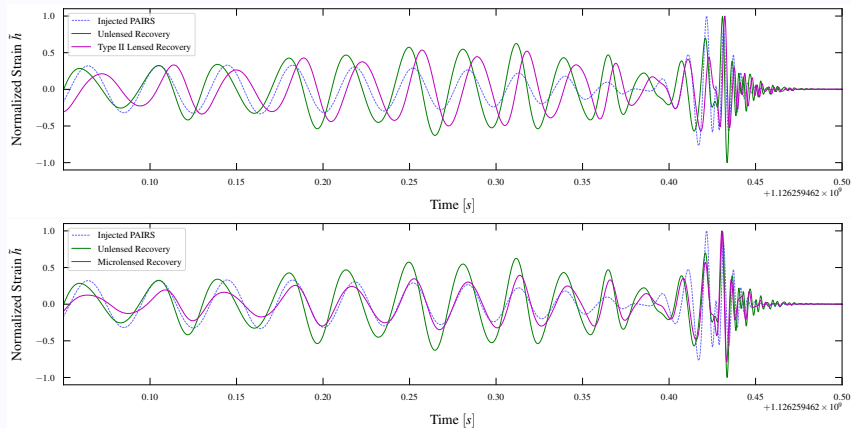
Type II Lensing and Microlensing Effects

Single-Image Lensing Analysis



- **Strong Lensing (Type II):** It is an effect of scaling the signal's amplitude, a temporal delay and an overall phase shift. Resulting in non-trivial properties in the time domain for certain parameters.
- **Microlensing:** Exhibits frequency-dependent amplification due to wave-optics effects. Resemblance to the beating patterns in overlapping signals.

Resemblance with Overlapping transients



Recovery of overlapping transient with Type II Lensing (*Top*) and Microlensing (*Bottom*) effects.

Study Overview

- Systematically vary key parameters influencing waveform evolution:
 - Chirp mass ratio: $\mathcal{M}_B/\mathcal{M}_A$
 - SNR ratio: $\text{SNR}_B/\text{SNR}_A$
 - Coalescence time difference: Δt_c
- Recoveries for **Unlensed** (single waveform), **Type II Lensed** and **Microlensed**.

Individual Case	Population Study
• $\mathcal{M}_B/\mathcal{M}_A \in \{0.5, 1, 2\}$ • $\text{SNR}_B/\text{SNR}_A \in \{0.5, 1\}$ • $\Delta t_c \in [-1, 1] \text{ s}$	• $\mathcal{M}_B/\mathcal{M}_A \in [0.1, 10]$ • $\text{SNR}_B/\text{SNR}_A \in [0.1, 10]$ • $\Delta t_c \in [-1, 1] \text{ s}$
Parameter estimation (for 60 signals).	Fitting factor ($\sim \mathcal{O}(5000)$ signals).

Analysis Techniques

- **Parameter Estimation:** Bayesian inference [Veitch, J. and Vecchio, A., 2010]:

$$\mathcal{L}(d|\theta) \propto \exp \left[-\frac{1}{2} \langle d - h(\theta) | d - h(\theta) \rangle \right]$$

$$\underbrace{\log_{10} \mathcal{B}_U^L}_{\text{Bayes Factor}} = \log_{10} \mathcal{Z}_L - \log_{10} \mathcal{Z}_U, \quad \mathcal{Z}_M = \int d\theta \mathcal{L}(d|\theta, \mathcal{H}_M) \pi_M(\theta|\mathcal{H}_M)$$

- **Fitting Factor:** Maximizing waveform overlap [Owen, B.J., 1996]:

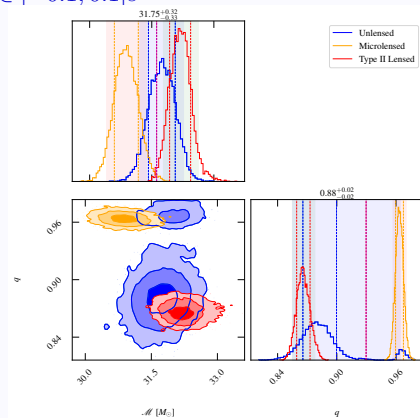
$$\mathcal{M}[h_1, h_2] = \max_{t_c, \Phi_c} \frac{\langle h_1 | h_2 \rangle}{\sqrt{\langle h_1 | h_1 \rangle \langle h_2 | h_2 \rangle}}.$$

$$\mathcal{F} = \max_{\lambda} \mathcal{M}[h_1, h_2(\lambda)], \quad \log_{10} \mathcal{B}_U^L = (\mathcal{F}_L^2 - \mathcal{F}_U^2) \frac{\text{SNR}^2}{2}$$

Base Results: Unlensed Parameter Estimation Recoveries

$\mathcal{M}_B/\mathcal{M}_A \in \{0.5, 1, 2\}$, $\text{SNR}_B/\text{SNR}_A \in \{0.5, 1\}$, $\Delta t_c \in [-0.1, 0.1]s$

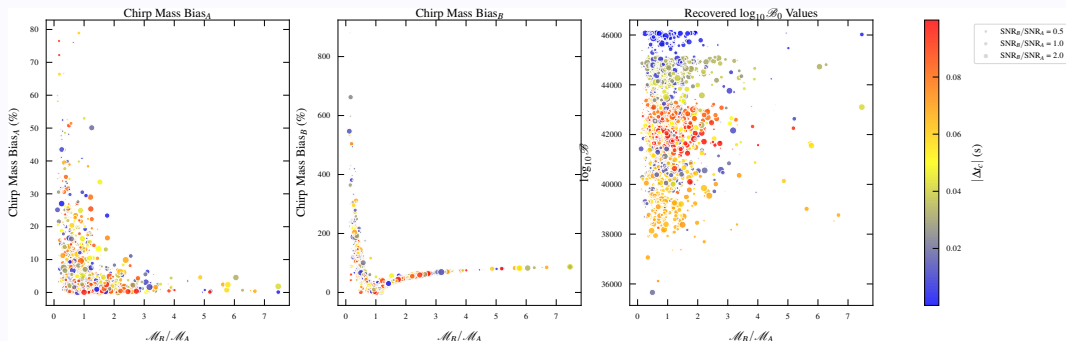
- Overlap induces significant parameter biases (e.g., in chirp mass).
- Higher-SNR signals dominate recovery.
- Bimodal distributions emerge in strongly overlapping cases.



$\mathcal{M}_B/\mathcal{M}_A = 2$, $\text{SNR}_B/\text{SNR}_A = 1$, $\Delta t_c = 0.01$ s.

Population Fitting Factor Results: Unlensed Signals

- Confirms the individual case dependencies and biases.
- Evident biases increase at stronger overlaps and comparable chirp mass, SNRs.

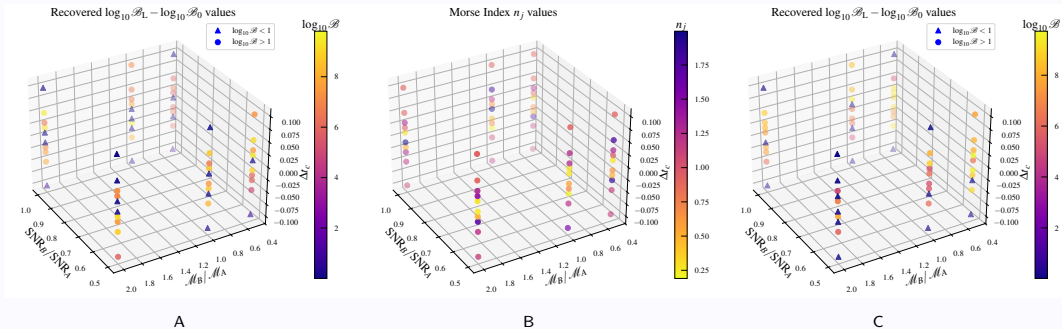


Type II Lensed Parameter Estimation Recoveries

$\mathcal{M}_B/\mathcal{M}_A \in \{0.5, 1, 2\}$, $\text{SNR}_B/\text{SNR}_A \in \{0.5, 1\}$, $\Delta t_c \in [-0.1, 0.1]s$

A : Fixed Morse phase shows distinct Bayes factor differences over the unlensed case.

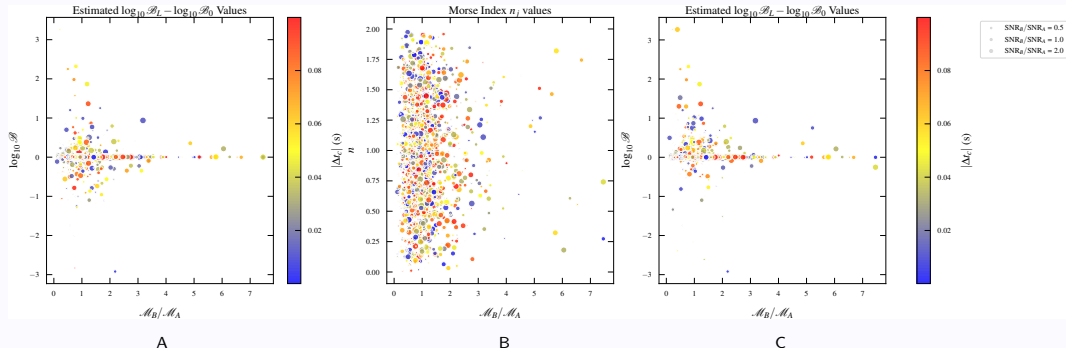
B,C : Allowing the Morse phase to vary improves lensing characterization.



Population Fitting Factor Results: Type II Lensed Signals

A : Fixed Morse phase presents mild support for Type II lensing recovery, except at comparable chirp mass, SNRs.

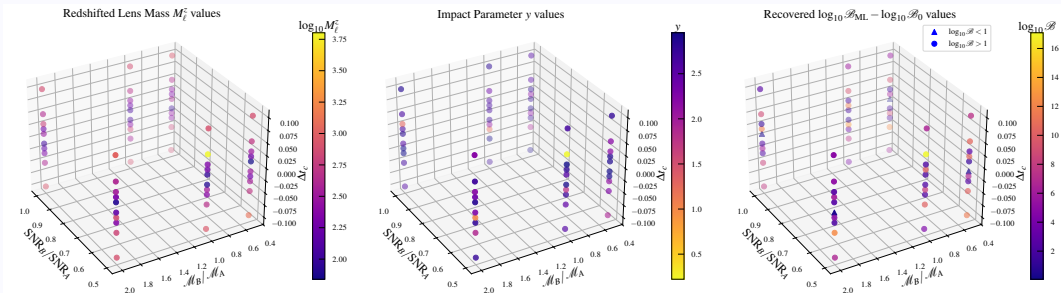
B, C : Allowing the Morse phase to vary shows greater support for strong lensing.



Microlensed Parameter Estimation Recoveries

$\mathcal{M}_B/\mathcal{M}_A \in \{0.5, 1, 2\}$, $\text{SNR}_B/\text{SNR}_A \in \{0.5, 1\}$, $\Delta t_c \in [-0.1, 0.1]s$

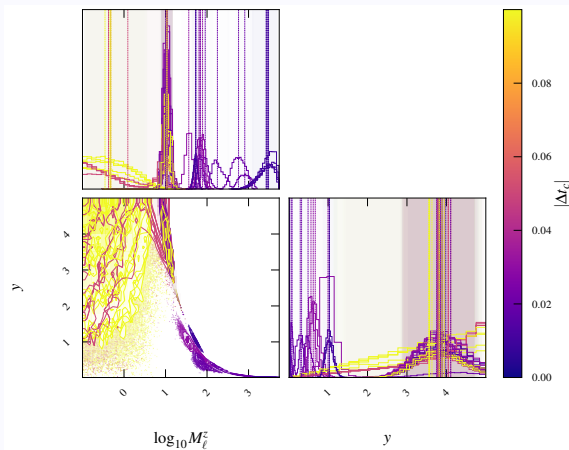
- Microlensed templates yield stronger support for strongly overlapping signals.
- Recovered lens parameters (M_ℓ^z , y) show dependencies with Δt_c .



Gravitational Wave Microlensing Parameter Estimation Posteriors

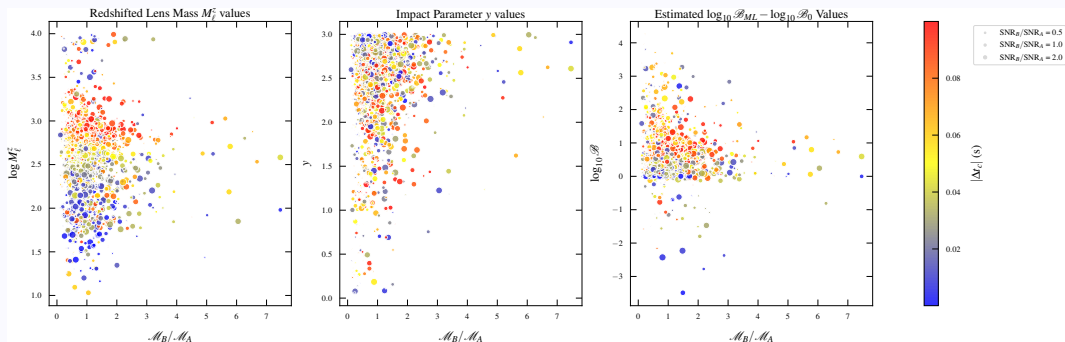
$$\mathcal{M}_B/\mathcal{M}_A \in \{0.5, 1, 2\}, \text{ SNR}_B/\text{SNR}_A \in \{0.5, 1\}, \Delta t_c \in [-0.1, 0.1]s$$

- For cases with significant overlap, the microlensed recovery tends to favour **larger** lens masses and **smaller** impact parameters.
- Decrease in support for lensing at higher difference in coalescence times.



Population Fitting Factor Results: Microlensed Signals

- Significant support for microlensing at specific parameter regimes.
- Larger lens mass and smaller impact parameters recovered at significant chirp mass and SNR ratios.

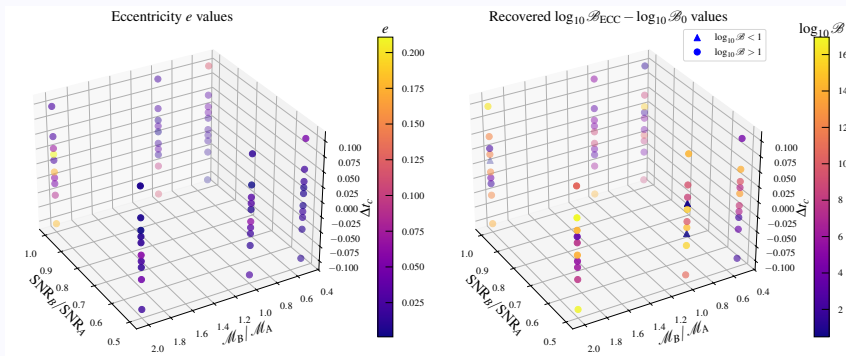


Conclusions

- Overlapping signals lead to significant biases in single signal unlensed parameter recovery.
- Further, we find that Type II lensing and microlensing signals can mimic overlapping effects, especially in the strong overlap regime.
- Advanced parameter estimation methods are essential to disentangle these effects.

Further Work

- Reanalyzing microlensing signals with the rectified Fourier convention.
- Future work will explore eccentric waveforms mimicking as a potential follow-up.



References

-  Samajdar, A. et al., 2021. *Biases in parameter estimation from overlapping gravitational-wave signals in the third-generation detector era*. Physical Review D, 104(4), p.044003.
-  Relton, P. and Raymond, V., 2021. *Parameter estimation bias from overlapping binary black hole events in second generation interferometers*. Physical Review D, 104(8), p.084039.
-  Janquart, J. et al., 2023. *Analyses of overlapping gravitational wave signals using hierarchical subtraction and joint parameter estimation*. Monthly Notices of the Royal Astronomical Society, 523(2), pp.1699-1710.
-  Veitch, J. and Vecchio, A., 2010. *Bayesian coherent analysis of in-spiral gravitational wave signals with a detector network*. Physical Review D—Particles, Fields, Gravitation, and Cosmology, 81(6), p.062003.
-  Owen, B.J., 1996. *Search templates for gravitational waves from inspiraling binaries: Choice of template spacing*. Physical Review D, 53(12), p.6749.

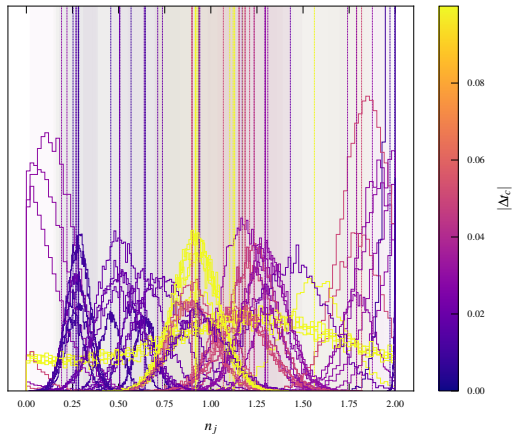
The background of the slide features a complex gravitational lensing effect. It shows a bright, yellowish-white light source at the bottom center, which is being distorted by the gravity of a massive object (likely a black hole or galaxy) positioned slightly to the left. This creates a series of concentric, wavy lines of light and shadow, resembling ripples in spacetime. The colors transition from a pale blue at the top to a warm yellow and orange near the light source, with some darker, shadowed regions in the upper left.

Thank You!

Questions? Comments?

Type II Lensing Parameter Estimation Posteriors

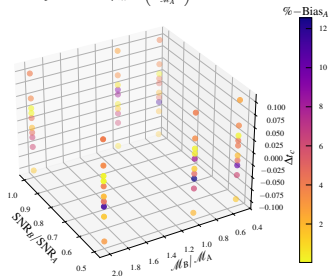
$\mathcal{M}_B/\mathcal{M}_A \in \{0.5, 1, 2\}$, $\text{SNR}_B/\text{SNR}_A \in \{0.5, 1\}$, $\Delta t_c \in [-0.1, 0.1]s$



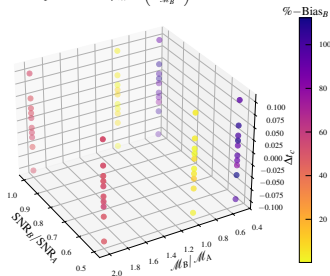
Unlensed Parameter Estimation Results

$\mathcal{M}_B/\mathcal{M}_A \in \{0.5, 1, 2\}$, $\text{SNR}_B/\text{SNR}_A \in \{0.5, 1\}$, $\Delta t_c \in [-0.1, 0.1]s$

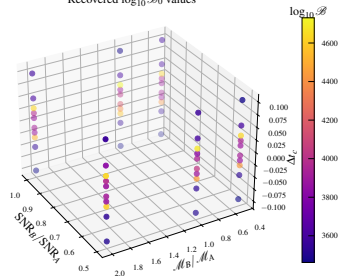
Chirp Mass Bias_A $\beta_{\mathcal{M}} = \left(\frac{\mathcal{M} - \mathcal{M}_A}{\mathcal{M}_A} \right) \% \text{ values}$



Chirp Mass Bias_B $\beta_{\mathcal{M}} = \left(\frac{\mathcal{M} - \mathcal{M}_B}{\mathcal{M}_B} \right) \% \text{ values}$



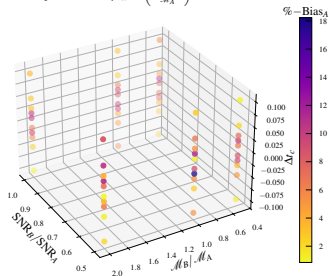
Recovered $\log_{10} \mathcal{P}_0$ values



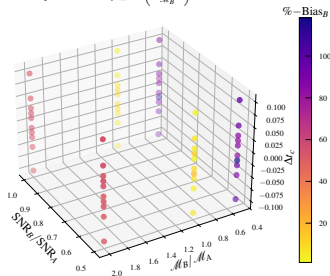
Unlensed Fitting Factor Results

$\mathcal{M}_B/\mathcal{M}_A \in \{0.5, 1, 2\}$, $\text{SNR}_B/\text{SNR}_A \in \{0.5, 1\}$, $\Delta t_c \in [-0.1, 0.1]s$

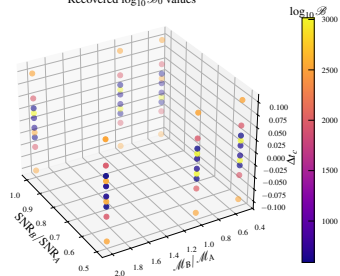
Chirp Mass Bias_A $\beta_{\mathcal{M}} = \left(\frac{\mathcal{M} - \mathcal{M}_A}{\mathcal{M}_A} \right) \% \text{ values}$



Chirp Mass Bias_B $\beta_{\mathcal{M}} = \left(\frac{\mathcal{M} - \mathcal{M}_B}{\mathcal{M}_B} \right) \% \text{ values}$

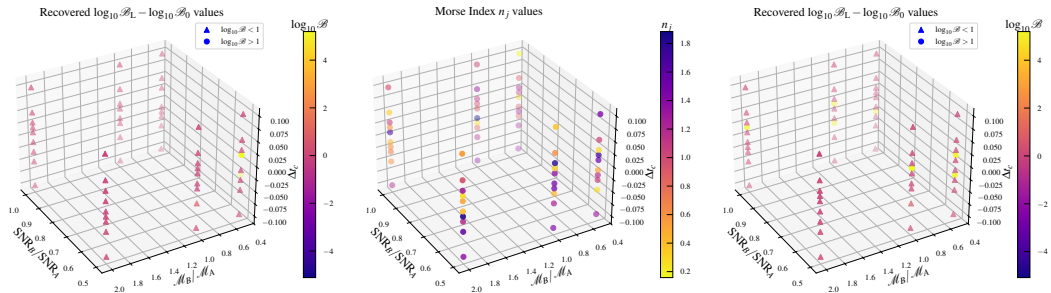


Recovered $\log_{10} \mathcal{P}_0$ values



Type II Lensed Fitting Factor Results

$\mathcal{M}_B/\mathcal{M}_A \in \{0.5, 1, 2\}$, $\text{SNR}_B/\text{SNR}_A \in \{0.5, 1\}$, $\Delta t_c \in [-0.1, 0.1]\text{s}$



Microlensed Fitting Factor Results

$\mathcal{M}_B/\mathcal{M}_A \in \{0.5, 1, 2\}$, $\text{SNR}_B/\text{SNR}_A \in \{0.5, 1\}$, $\Delta t_c \in [-0.1, 0.1]s$

